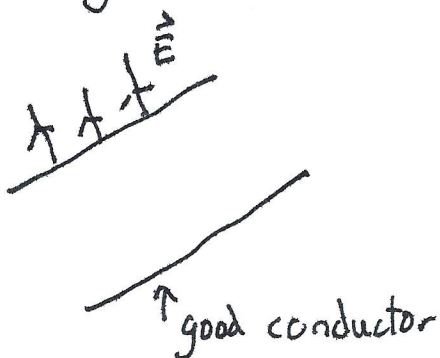


Phys 110B 06 Nov 20

"Antenna Paradox"

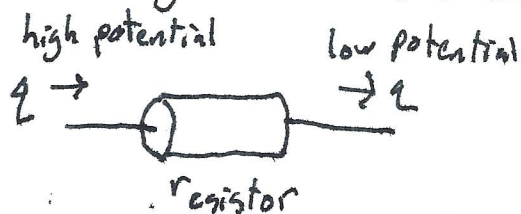
use a good conductor for efficiency



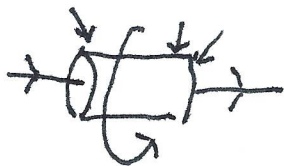
$$\hat{n} \cdot \vec{S} \Big|_{\substack{\text{surface} \\ \text{of conductor}}} \propto (\vec{E} \times \vec{B}) \cdot \hat{n} \propto (\hat{n} \times \vec{E}) \cdot \vec{B} = 0$$

No Poynting flux out of the antenna!

Antenna is more like a guide of radiation



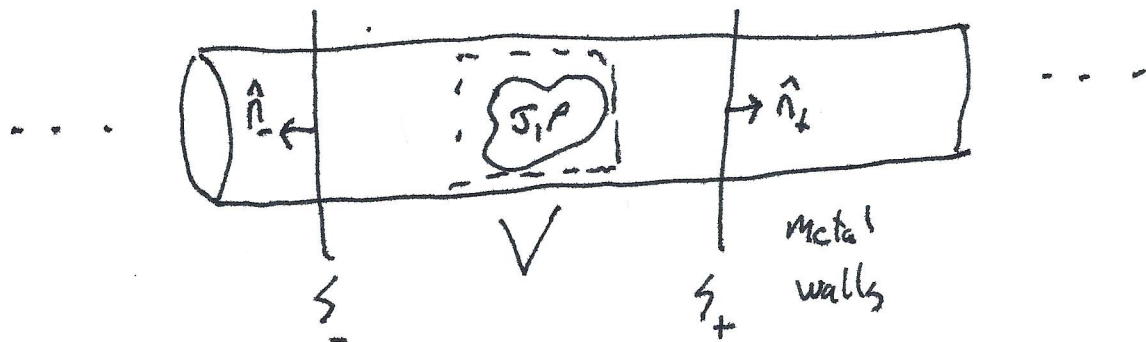
fields



\vec{S} is radial in towards the resistor

Possibly taking our book-keeping too seriously

Radiation by a localized source into a waveguide



Localized $\vec{J} \sim e^{-i\omega t}$ prescribed source

+ right moving waves

- left moving waves

$$\vec{E}_j^+(\vec{x}; \omega) = \left(\vec{E}_j(x, y; i\omega) + \hat{z} E_{jz}(x, y; i\omega) \right) e^{ik_j z}$$

real transverse profile

$\rightarrow \hat{z} E_{jz}(x, y; i\omega)$

$$\omega = \omega_j(k_j)$$

$$\vec{H}_j^+(\vec{x}; \omega) = (\vec{H}_j + \hat{z} H_{jz}) e^{ik_j z}$$

$$\vec{E}_j^-(\vec{x}; \omega) = (\vec{E}_j - \hat{z} E_{jz}) e^{-ik_j z}$$

$$\vec{H}_j^-(\vec{x}; \omega) = (-\vec{H}_j + \hat{z} H_{jz}) e^{-ik_j z}$$

Can also choose $\frac{1}{2} \int da \hat{z} \cdot (\vec{E}_j \times \vec{H}_l) = \delta_{jl} \frac{1}{2Z_j}$

wave impedance

"reciprocity" relationships

$$\vec{\nabla} \cdot (\vec{E} \times \vec{H}_j^\pm - \vec{E}_j^\pm \times \vec{H})$$

↑
actual fields

↑
wave guide modes

in presence of the sources

$$\vec{\nabla} \cdot (\vec{E} \times \vec{H}_j^\pm - \vec{E}_j^\pm \times \vec{H})$$

$$\begin{aligned} \rightarrow \vec{H}_j^\pm \cdot (\vec{\nabla} \times \vec{E}) - \vec{E} \cdot (\vec{\nabla} \times \vec{H}_j^\pm) - \vec{H} \cdot (\vec{\nabla} \times \vec{E}_j^\pm) + \vec{E}_j^\pm \cdot (\vec{\nabla} \times \vec{H}) \\ \parallel \quad \parallel \quad \parallel \quad \parallel \\ i\omega \vec{B} \quad \vec{0} + \epsilon_0 i\omega \vec{E}_j^\pm \quad i\omega \vec{B}_j^\pm \quad \vec{J} + \epsilon_0 i\omega \vec{E} \\ \parallel \quad \parallel \\ \mu \vec{H} \end{aligned}$$

$$= \dots = \vec{J} \cdot \vec{E}_j^\pm \quad \text{Lorentz reciprocity formula}$$

integrate over region V

$$\int_{s_- + s_+ + s_w} (\vec{E} \times \vec{H}_j^\pm - \vec{E}_j^\pm \times \vec{H}) \cdot \hat{n} da = \int_V d^3x \vec{J} \cdot \vec{E}_j^\pm$$

$$\vec{E}_j^\pm|_{s_w} \propto \hat{n}|_{s_w}$$

no contribution on left from the walls s_w

$$\vec{E}|_{s_w} \propto \hat{n}|_{s_w}$$

$$\Delta n \ s_+ \quad \vec{E}_\perp = \sum_j a_j^+ \vec{E}_j^+ e^{ik_j z} \quad \hat{n}|_{s_+} \propto \hat{z}$$

↑ expansion coefficients

$$\Delta n \ s_- \quad \vec{E}_\perp = \sum_j a_j^- \vec{E}_j^- e^{-ik_j z} \quad \hat{n}|_{s_-} \propto -\hat{z}$$

$$\int_V \vec{J} \cdot \vec{E}_j^+ d^3x' = \sum_j a_j^+ \int_{s_+} (\vec{E}_j^+ \times \vec{H}_\ell^- - \vec{E}_\ell^- \times \vec{H}_j^+) \cdot \hat{z} da$$

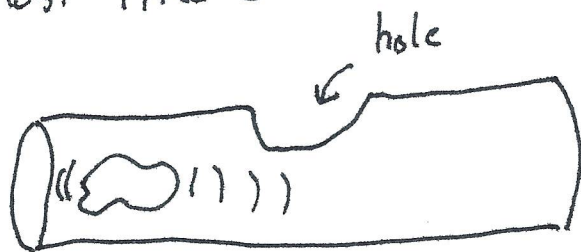
$$- \sum_j a_j^- \int_{s_-} (\vec{E}_j^- \times \vec{H}_\ell^- - \vec{E}_\ell^- \times \vec{H}_j^-) \cdot \hat{z} da \rightarrow 0 \text{ due to orthogonality}$$

$$\int_V \vec{J} \cdot \vec{E}_j^- d^3x' = \sum_j a_j^+ \int_{s_+} (\vec{E}_j^+ \times \vec{H}_\ell^- - \vec{E}_\ell^- \times \vec{H}_j^+) \cdot \hat{z} da = *$$

$$\Rightarrow a_{\ell}^{\dagger} = -\frac{Z_0}{2} \int_V \vec{J} \cdot \vec{E}_{\ell}^{\dagger} d^3x \quad \text{"Joule" work integral}$$

← overlap integral
Over the source

Current density \vec{J} tends to radiate into the modes that "look" most like \vec{J}



Radiation from relativistic sources (mostly into vacuum)

"radiation" fields

$$\vec{E}_{\text{rad}} = \frac{q}{4\pi\epsilon_0 c^2} \left[\frac{\hat{R} \times (\hat{R} - \vec{\beta}) \times \dot{\vec{u}}}{(R - \vec{\beta} \cdot \hat{R})^3} \right]$$

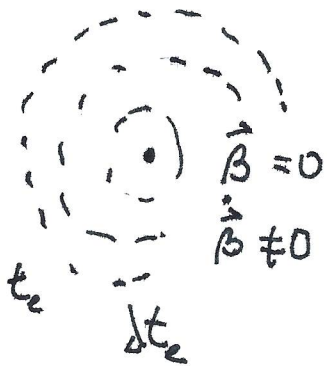
$$\vec{B}_{\text{rad}} = \left[\frac{\hat{R} \times \vec{E}_{\text{rad}}}{c} \right]$$

$$\Rightarrow \vec{E}_{\text{rad}} = \frac{q}{4\pi\epsilon_0} \left[\frac{\hat{R} \times (\hat{R} - \vec{\beta}) \times \dot{\vec{\beta}}}{(1 - \vec{\beta} \cdot \hat{R})^3 c R} \right]_{\text{ret}}$$

$$\vec{B}_{\text{rad}} = \left[\frac{\hat{R} \times \vec{E}_{\text{rad}}}{c} \right]$$

$$\vec{\zeta}_{\text{rad}} = \frac{1}{\mu_0 c} |\vec{E}_{\text{rad}}|^2 \hat{R} \quad \text{far-field}$$

for a charge instantaneously at rest



look over some large sphere at some suitable later time

$$\frac{dP}{d\Omega} = R^2 \zeta_{\text{rad}} \cdot \hat{R} = \frac{\mu_0 q^2 a^2}{16\pi^2 c} \sin^2 \theta$$

$$\vec{a} = \frac{\dot{\vec{v}}}{R} = \text{acceleration of charge}$$

