

Phys 110B 23 Oct 20

Jefimenko Expressions

$$\vec{E}(\vec{x}, t) = \frac{1}{4\pi\epsilon_0} \int d^3x' \left(\frac{\hat{R}}{R^2} [\rho(\vec{x}', t')] \right]_{ret} + \frac{\hat{R}}{cR} \left[\frac{\partial}{\partial t'} \rho(\vec{x}', t') \right]_{ret} - \frac{1}{c^2} \frac{1}{R} \left[\frac{\partial}{\partial t'} \vec{J}(\vec{x}', t') \right]_{ret} \quad \begin{matrix} \vec{R} = \vec{x} - \vec{x}' & R = |\vec{R}| \\ \hat{R} = \frac{\vec{R}}{R} \end{matrix}$$

$$\vec{B}(\vec{x}, t) = \frac{\mu_0}{4\pi} \int d^3x' \left(\left[\vec{J}(\vec{x}', t') \right]_{ret} \times \frac{\hat{R}}{R^2} + \left[\frac{\partial}{\partial t'} \vec{J}(\vec{x}', t') \right]_{ret} \times \frac{\hat{R}}{cR} \right)$$

Panofsky - Phillips

K. MacDonald

$$\vec{E}(\vec{x}, t) = \frac{1}{4\pi\epsilon_0} \int d^3x' \frac{[\rho]_{ret} \hat{R}}{R^2} + \frac{1}{4\pi\epsilon_0} \int d^3x' \left(\frac{([\vec{J}]_{ret} \cdot \hat{R}) \hat{R} + ([\vec{J}]_{ret} \times \hat{R}) \times \hat{R}}{cR^2} + \frac{([\frac{\partial}{\partial t'} \vec{J}]_{ret} \times \hat{R}) \times \hat{R}}{c^2 R} \right)$$

← currents produce electric fields?

vanishes in steady state

$$\frac{\partial}{\partial t} \vec{J}(\vec{x}, t) = \vec{0}$$

$$\vec{\nabla} \cdot \vec{J}(\vec{x}, t) = 0$$

Fields from point charges

$$\rho(\vec{x}, t) = q \delta(\vec{x} - \vec{r}(t)) \quad \vec{j}(\vec{x}, t) = \hat{v}(t) \rho(\vec{x}, t)$$

with trajectory $\vec{r}(t)$

$$\Rightarrow \vec{E} = \frac{q}{4\pi\epsilon_0} \left(\left[\frac{\hat{R}}{R^2} \right]_{\text{ret}} + \frac{\partial}{\partial t} \left[\frac{\hat{R}}{KR} \right]_{\text{ret}} - \frac{\partial}{c^2 \partial t} \left[\frac{\hat{v}}{KR} \right]_{\text{ret}} \right)$$

$$\vec{B} = \frac{\mu_0 q}{4\pi} \left(\left[\frac{\hat{v} \times \hat{R}}{KR^2} \right]_{\text{ret}} + \frac{\partial}{c \partial t} \left[\frac{\hat{v} \times \hat{R}}{KR} \right]_{\text{ret}} \right)$$

$\hat{v} = \dot{\vec{r}}$ velocity of charge

$$|\hat{R}| = |\vec{x} - \vec{r}(t_r)| \quad t_r = t - \frac{|\vec{x} - \vec{r}(t_r)|}{c}$$

$$K = \text{"time compression factor"} = 1 - \frac{\hat{v} \cdot \hat{R}}{c} = 1 - \beta \cdot \hat{R}$$

$\Rightarrow \dots \Rightarrow$ Feynman-Heaviside Expressions for fields of a moving point charge

$$\vec{E} = \frac{q}{4\pi\epsilon_0} \left(\left[\frac{\hat{R}}{R^2} \right]_{\text{ret}} + \frac{[R]_{\text{ret}}}{c} \frac{\partial}{\partial t} \left[\frac{\hat{R}}{R^2} \right]_{\text{ret}} + \frac{\partial^2}{c^2 \partial t^2} \left[\hat{R} \right]_{\text{ret}} \right)$$

$$\vec{B} = \frac{\mu_0 q}{4\pi} \left(\left[\frac{\hat{v} \times \hat{R}}{K^2 R^2} \right]_{\text{ret}} + \frac{1}{c [R]_{\text{ret}}} \frac{\partial}{\partial t} \left[\frac{\hat{v} \times \hat{R}}{K} \right]_{\text{ret}} \right)$$

\vec{E} First term is Coulomb's Law evaluated at retarded time

Second term is like a 1st order term in Taylor series
correction for motion

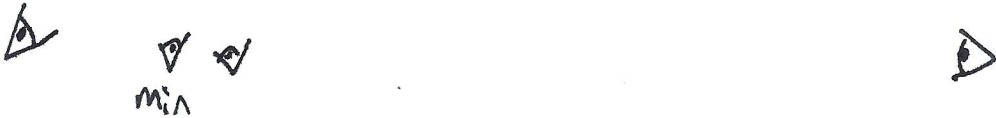
Third term

$$\frac{q}{4\pi\epsilon_0} \frac{\partial^2}{c^2 \partial t^2} \left[\hat{R} \right]_{\text{ret}} = \text{Radiation}$$

radiation is associated with
the acceleration of charges

Radiation is Nature's way of communicating changes in the state of motion of sources to remote locations.

oscillating electron



Different angles of observer see a projection

Magnetic radiation term

radiation must locally look like a plane wave

$$\vec{B} = \frac{[\hat{R} \times \dot{\vec{E}}]_{\text{ret}}}{c}$$

Liénard - Wiechert Potentials

Lorentz-Lorenz gauge potentials due to a point charge

$$\phi = \frac{q}{4\pi\epsilon_0} \left[\frac{1}{R\kappa} \right]_{\text{ret}} \quad \kappa = 1 - \hat{n} \cdot \vec{\beta}$$

τ of the particle

$$\vec{A} = \frac{\mu_0 q}{4\pi} \left[\frac{\dot{\vec{B}}}{cR\kappa} \right]_{\text{ret}}$$

$$\vec{E} = -\nabla\phi - \frac{\partial \vec{A}}{\partial t} = \dots = \frac{q}{4\pi\epsilon_0} \left[\frac{(1-\beta^2)(\hat{R}-\vec{\beta})}{R^2(1-\hat{n}\cdot\vec{\beta})^3} \right]_{\text{ret}}^{\#1} + \frac{q}{4\pi\epsilon_0} \left[\frac{\hat{R} \times ((\hat{R}-\vec{\beta}) \times \dot{\vec{\beta}})}{cR(1-\hat{n}\cdot\vec{\beta})^3} \right]_{\text{ret}}^{\#2}$$

$$\vec{B} = \dots = \frac{[\hat{R} \times \dot{\vec{E}}]_{\text{ret}}}{c}$$

$$\vec{B} = \dots = \left[\frac{\hat{R} \times \dot{\vec{E}}}{c} \right]_{\text{ret}}$$

*1 falls off like $\frac{1}{R^2}$ depends on $\vec{\beta}$ but not $\dot{\vec{\beta}}$ "velocity fields" non-radiative

*2 falls off like $\frac{1}{R}$ depends on $\dot{\vec{\beta}}$ "acceleration fields" radiative